## Extending the fluid dynamic description of heavy-ion collisions to times before the collision

by Andreas Kirchner

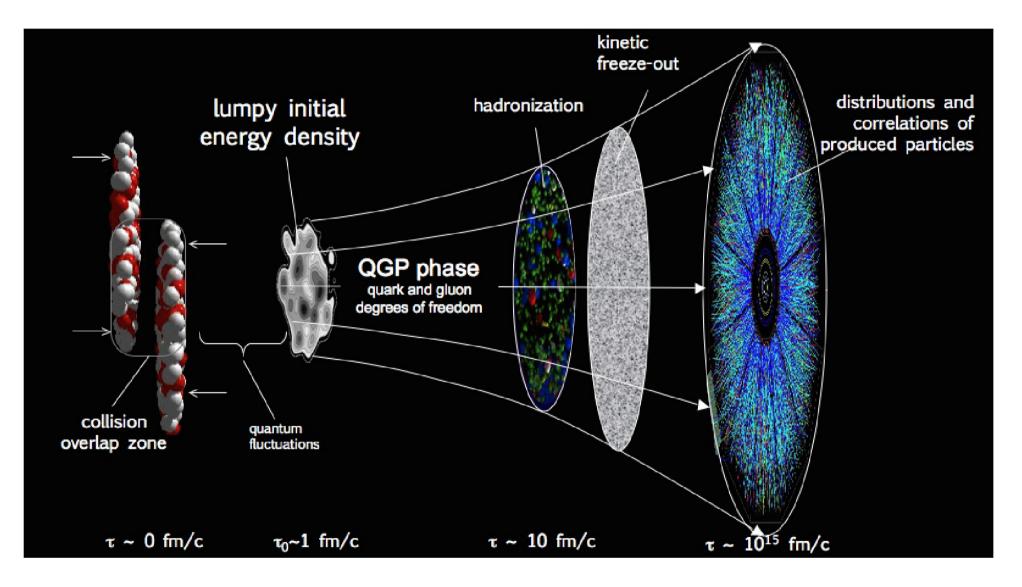
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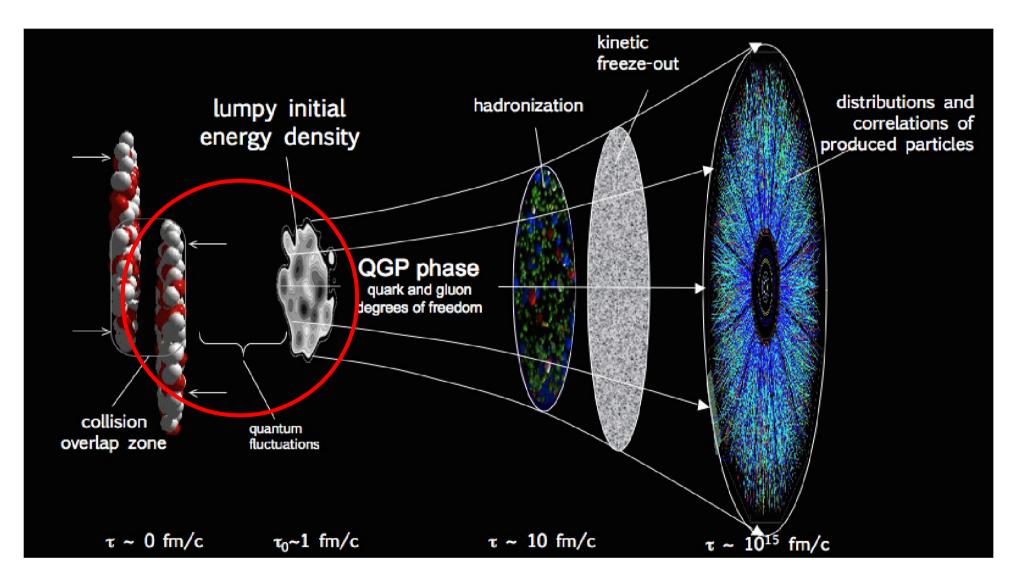




## Standard model of heavy-ion collisions



# Standard model of heavy-ion collisions



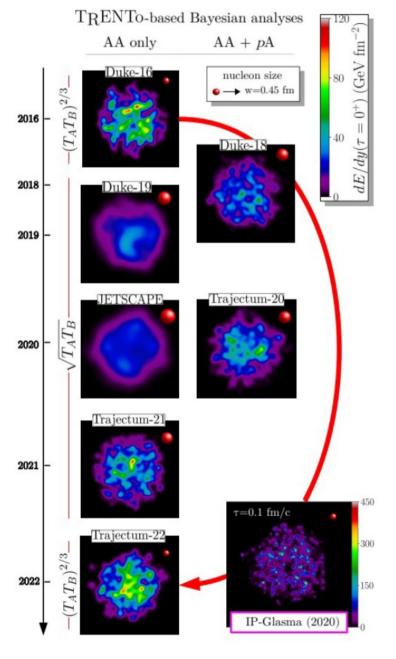
# Current model of initial conditions

- Different "models on market" for initializing hydro:
  - TrenTo
  - IP-Glasma
  - Color Glass Condensate
- TrenTo checks geometrically for collisions of nucleons  $\rightarrow$  Reduced thickness functions  $T_{A/B}$

$$\frac{dE}{dy} \propto \left(\frac{T_A^p + T_b^p}{2}\right)^{q/p}$$

What are the optimal values for p and q ?

# Little history of initial conditions



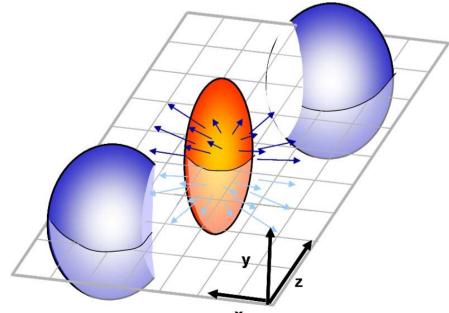
- Values of p and q determined by bayesian analysis, fit to data
  - $\rightarrow$  Values changed back and forth
  - $\rightarrow$  Interpretation of geometric picture changes between entropy and energy density
  - $\rightarrow$  Can this step in modelling be evaded?
    - Can hydro describe full collision?

2208.06839

# Working principle

- Cartesian coordinates (t, x, y, z)
- For now: only consider longitudinal expansion
- Fluid fields reduce to

$$\phi = (T, u, \pi^{zz}, \pi_{\text{Bulk}}, \nu, \mu)$$



# Hydrodynamic setting

 EoM derived from energy-momentum and baryon number current conservation

$$\nabla_{\mu}T^{\mu\nu} = 0 \qquad \qquad \nabla_{\mu}n^{\mu} = 0$$

$$u^{\mu}\partial_{\mu}\epsilon + (\epsilon + p + \pi_{\text{Bulk}})\nabla_{\mu}u^{\mu} + \pi^{\mu\nu}\nabla_{\mu}u_{\nu} = 0$$

$$(\epsilon + p + \pi_{\text{Bulk}})u^{\mu}\nabla_{\mu}u^{\nu} + \Delta^{\mu\nu}\partial_{\mu}(p + \pi_{\text{Bulk}}) + \Delta^{\nu}_{\alpha}\nabla_{\mu}\pi^{\mu\alpha} = 0$$

$$u^{\mu}\partial_{\mu}n + n\nabla_{\mu}u^{\mu} + \nabla_{\mu}\nu^{\mu} = 0$$

 Energy-mometum tensor is based on tensor decomposition with respect to time-like eigenvector (i.e. Fluid velocity)

$$T^{\mu\nu} = \epsilon u^{\mu} u^{\nu} + (p + \pi_{\text{Bulk}}) \Delta^{\mu\nu} + \pi^{\mu\nu}$$
$$n^{\mu} = n u^{\mu} + \nu^{\mu}$$

# Hydrodynamic setting

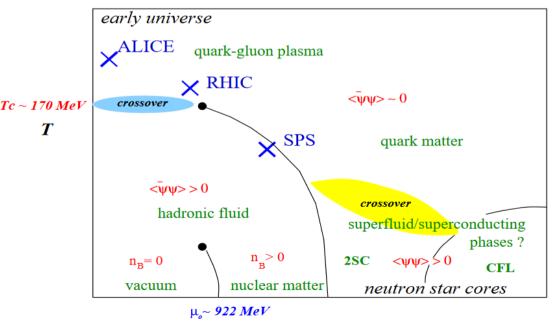
• Supplemental equations for  $\pi^{\mu\nu}$ ,  $\pi_{Bulk}$  and  $\nu^{\mu}$ : Second order Israel-Stewart and diffusion current equations

$$\begin{aligned} \tau_{H} \Delta^{\alpha}_{\beta} u^{\mu} \nabla_{\mu} \nu^{\beta} + \nu^{\alpha} + \kappa \left(\frac{nT}{\epsilon+p}\right)^{2} \Delta^{\alpha\beta} \partial_{\beta} \left(\frac{\mu_{B}}{T}\right) &= 0 \\ P^{\mu\nu\rho}_{\sigma} \left[\tau_{S} (u^{\lambda} \nabla_{\lambda} \pi^{\sigma}_{\rho} - 2\pi^{\sigma\lambda} \omega_{\rho\lambda}) + 2\eta \nabla_{\rho} u^{\sigma}\right] + \pi^{\mu\nu} &= 0 \\ \tau_{\text{Bulk}} u^{\mu} \partial_{\mu} \pi_{\text{Bulk}} + \pi_{\text{Bulk}} + \zeta \nabla_{\mu} u^{\mu} &= 0 \\ \text{Israel-Stewart} \qquad \text{Ideal Navier-Stokes} \end{aligned}$$

• Isreal-Stewart: Introduce relaxation time  $\mathcal{T}$   $\rightarrow$  Equations remain valid outside equilibrium

# Nuclear droplet model

- Idea of describing nucleus as liquid not new!
   → Bethe-Weizsäcker-formula (1935)
- Use fluid variables to describe nucleus
  - $\rightarrow$  Fluid system described by energy-momentum tensor
- Single nucleus sits at vacuum-nuclear matter phase transition  $\rightarrow T^{\mu\nu}$  describes nucleus and vacuum!



# Adding energy & momentum

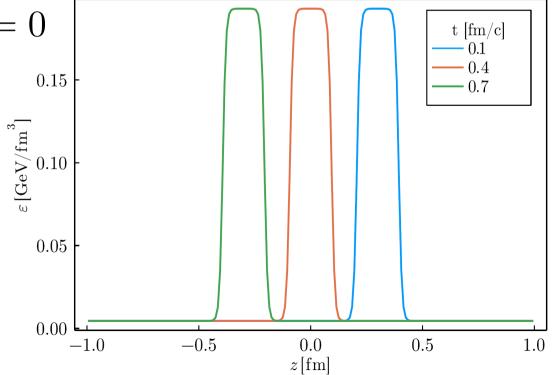
Incomming nuclei described by

$$T^{\mu\nu}_{\to/\leftarrow} = \epsilon_{\to/\leftarrow} u^{\mu}_{\to/\leftarrow} u^{\nu}_{\to/\leftarrow}$$

- Nuclei sit at nuclear phase transition
  - Initial energy density  $\epsilon = \mu_{crit} n$

 $- EoM simplify to <math>u^{\mu} \partial_{\mu} n = 0$ 

Free streaming nuclei



# Adding energy & momentum

 Energy-momentum tensor and number density current of collision system

$$T^{\mu\nu}_{\rm coll} = T^{\mu\nu}_{\rightarrow} + T^{\mu\nu}_{\leftarrow} \qquad \qquad n^{\mu}_{\rm coll} = n^{\mu}_{\rightarrow} + n^{\mu}_{\leftarrow}$$

Obtain fluid variables via Landau matching

$$T^{\mu}_{\nu}u^{\nu} = -\epsilon u^{\mu}, \quad u_{\mu}u^{\mu} = -1$$

• Number density and diffusion current given by

$$\begin{pmatrix} \gamma_{\rightarrow}n_{\rightarrow} + \gamma_{\leftarrow}n_{\leftarrow} \\ 0 \\ 0 \\ \gamma_{\rightarrow}\beta_{\rightarrow}n_{\rightarrow} + \gamma_{\leftarrow}\beta_{\leftarrow}n_{\leftarrow} \end{pmatrix} \stackrel{!}{=} \begin{pmatrix} \gamma n + \beta \nu \\ 0 \\ 0 \\ \gamma\beta n + \nu \end{pmatrix}$$

## Adding energy & momentum

• Viscous corrections given by

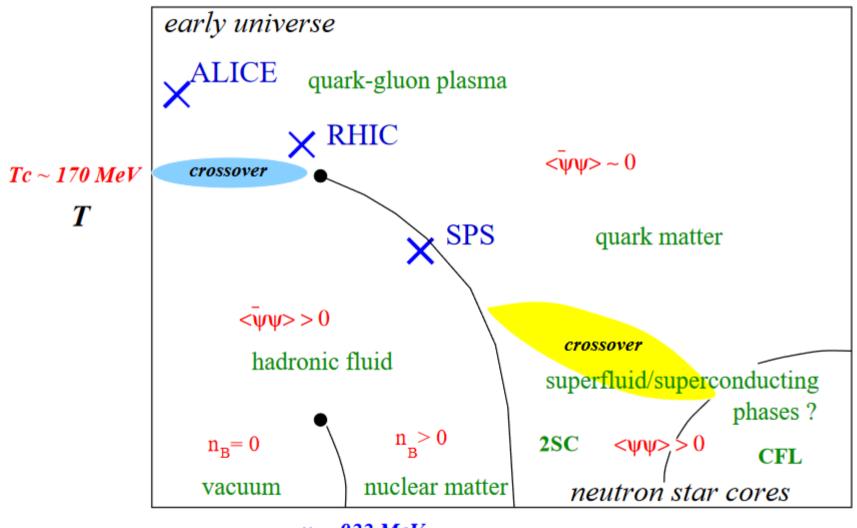
$$\pi_{\text{Bulk}} = \frac{1}{3} \Delta^{\mu\nu} T_{\mu\nu} - p(T,\mu) \quad \pi^{\mu\nu} = T^{\mu\nu} - \epsilon u^{\mu} u^{\nu} - (p + \pi_{\text{Bulk}}) \Delta^{\mu\nu}$$

Invert EoS to obtain full set of fluid variables

$$n(T,\mu) = n_{\text{coll}}$$
  
 $\epsilon(T,\mu) = \epsilon_{\text{coll}}$ 

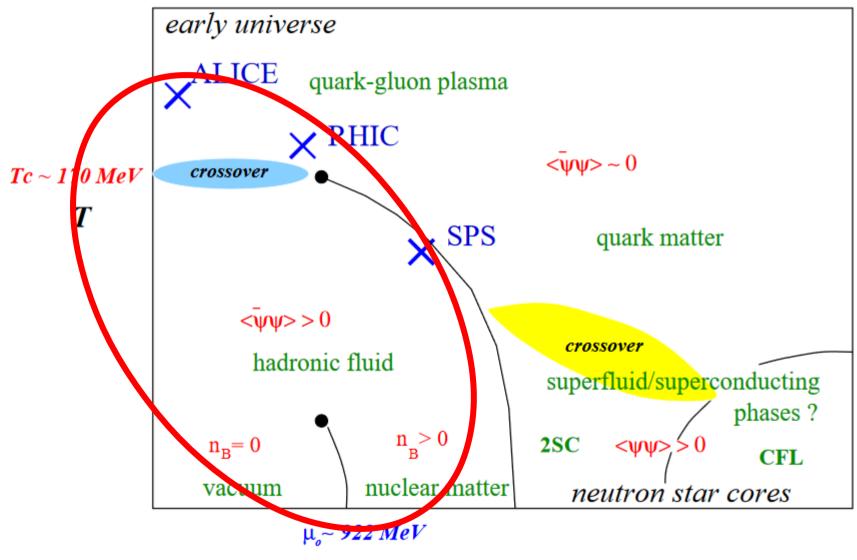
-Which EoS to use to cover large range in T and  $\mu$ ?

# Equation of state



μ<sub>o</sub>~ 922 MeV

## Equation of state



μ

#### EoS – High T

• At high T and low density  $\rightarrow$  Lattice QCD

$$p_{\text{LQCD}} = p(T) + T^4 \left(\frac{1}{2!}\chi_{2B}(T)\left(\frac{\mu}{T}\right)^2 + \frac{1}{4!}\chi_{4B}(T)\left(\frac{\mu}{T}\right)^4 \frac{1}{6!}\chi_{6B}(T)\left(\frac{\mu}{T}\right)^6\right)$$

early anivers

crossover

 $\langle \bar{\psi}\psi \rangle > 0$ 

hadronic fluid

 $Tc \sim 1$ 

quark-gluon plasma

 $n_{\rm B}^{>0}$ 

nuclea

SPS

hatter

μ

 $\langle \bar{\psi}\psi \rangle \sim 0$ 

crossover

2SC <\\\>

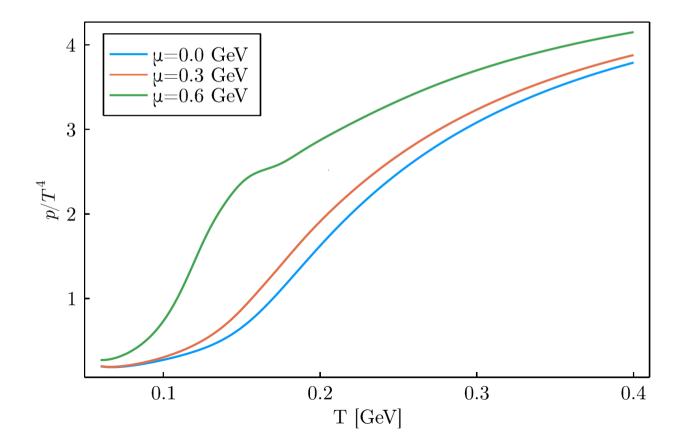
quark matter

superfluid/superconducting

neutron star cores

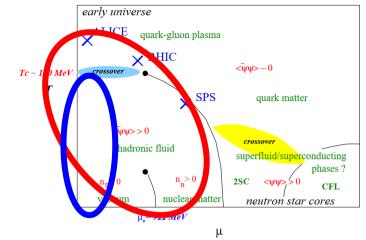
phases ?

CFL



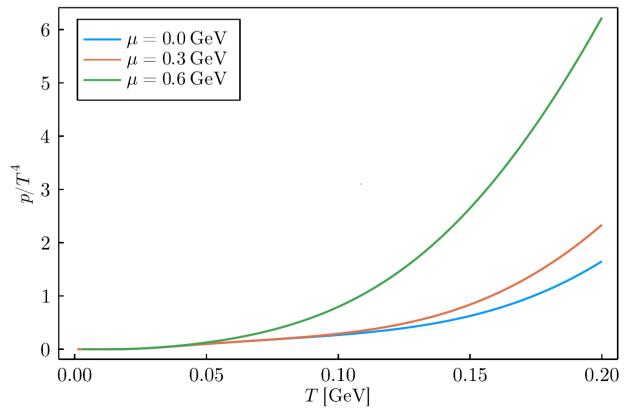
# EoS – Low T, low density

- Low T and low density  $\rightarrow$  Hadron Resonance Gas



$$p_{\mathrm{HRG}}(T,\mu) = \sum_{i} d_i p_{\mathrm{FG}}(T, B_i\mu; m_i) + \sum_{i} d_i p_{\mathrm{BG}}(T; m_i)$$

• Pressure given by sum of partial pressure of constituents



# EoS – Low T, high density

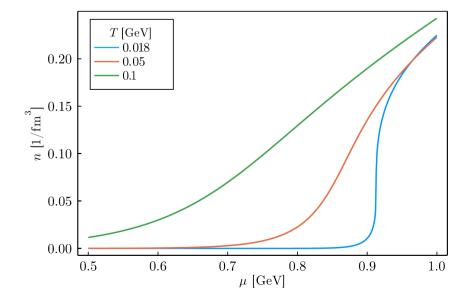
• Low T and high density  $\rightarrow$  Walecka model

- Effective model of protons and neutrons with omega and scalar meson exchange
- Pressure in mean-field approximation: [1202.1671]

 $p_{\rm WM}(T,\mu) = 4p_{\rm FG}(T,\mu^*;m_N^*) + 4p_{\rm FG}(T,-\mu^*;m_N^*) - \frac{1}{2}m_\sigma^2\bar{\sigma}^2 + \frac{1}{2}m_\omega^2\bar{\omega}_0^2$ 

· Mean-fields determined by gap equations

$$\bar{\omega}_{0} = \frac{g_{\omega}}{m_{\omega}^{2}} \frac{\partial P_{FD}}{\partial \mu^{*}}$$
$$\bar{\sigma} = -\frac{g_{\sigma}}{m_{\sigma}^{2}} \frac{\partial P_{FD}}{\partial m_{N}^{*}}$$



early universe

<ψψ>>0 hadronic fluid

quark-gluon plasma

 $\langle \bar{\Psi}\Psi \rangle \sim 0$ 

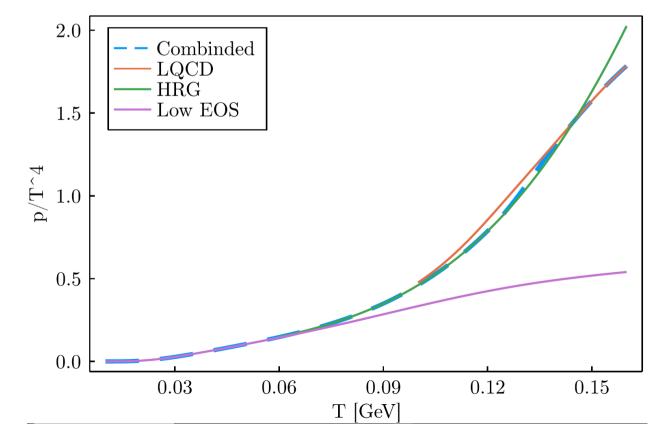
quark matter

nducting phases f

# Combining the EoS

- Low T: HRG w/o proton and neutron + Walecka
- High T: Transition from HRG to LQCD

$$p(T,\mu) = \Theta(T,\mu)p_{\text{LQCD}}(T,\mu) + (1 - \Theta(T,\mu))p_{\text{HRG}}(T,\mu)$$



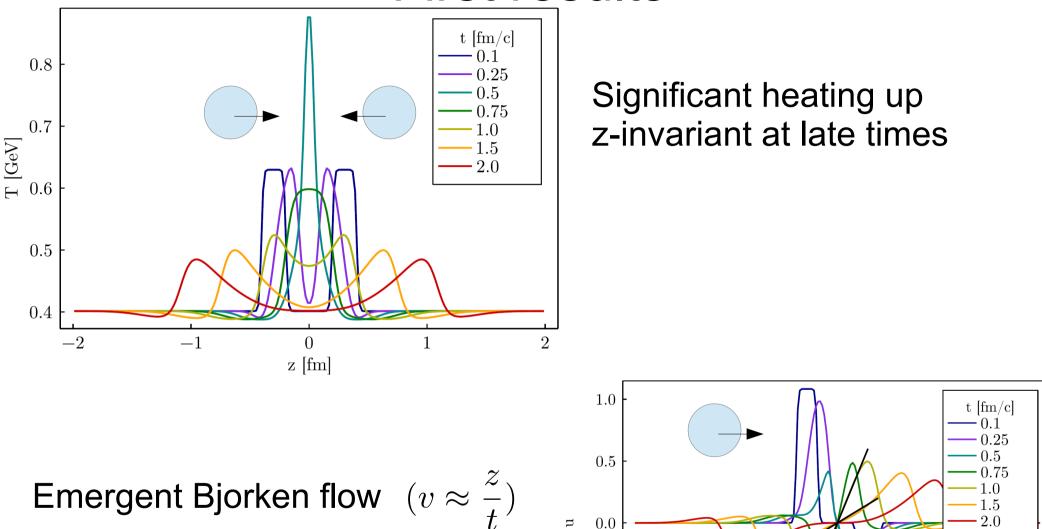
# Finite Temperature Landau matching

- EoS at low T tricky to handle (degenerate Fermi gas)  $p_{\rm HRG} \propto K_2(rac{m}{T}) \qquad p_{\rm Walecka} \propto {\rm Li}_{5/2}(-e^{rac{\mu-m}{T}})$
- Solution: Sommerfeld & asymptotic expansion
- For now: Do Landau matching at finite temperature

→Invert 
$$n(T_{BG}, \mu) = n_{drop}$$

 $\blacktriangleright$  Do Landau matching with  $T^{\mu\nu} = \epsilon(T_{BG}, \mu) u^{\mu} u^{\nu}$ 





-0.5

-1.0

-2

-1

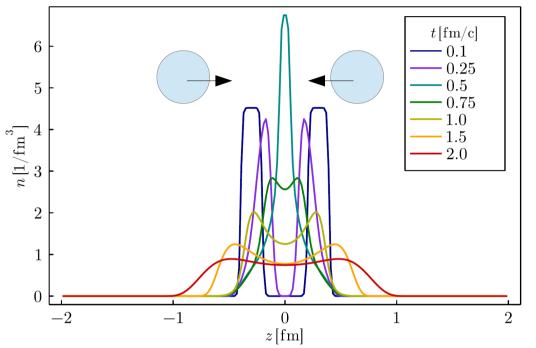
0

z [fm]

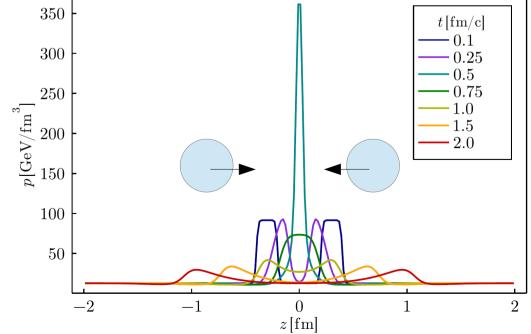
1

2

# First results

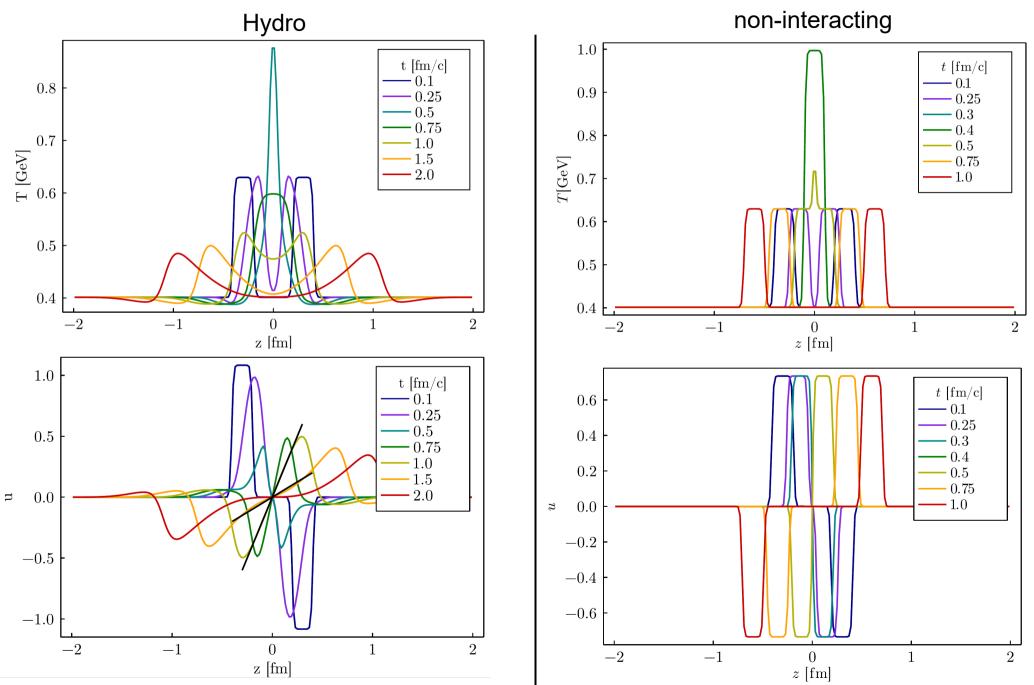


Matter clumps up in collision zone



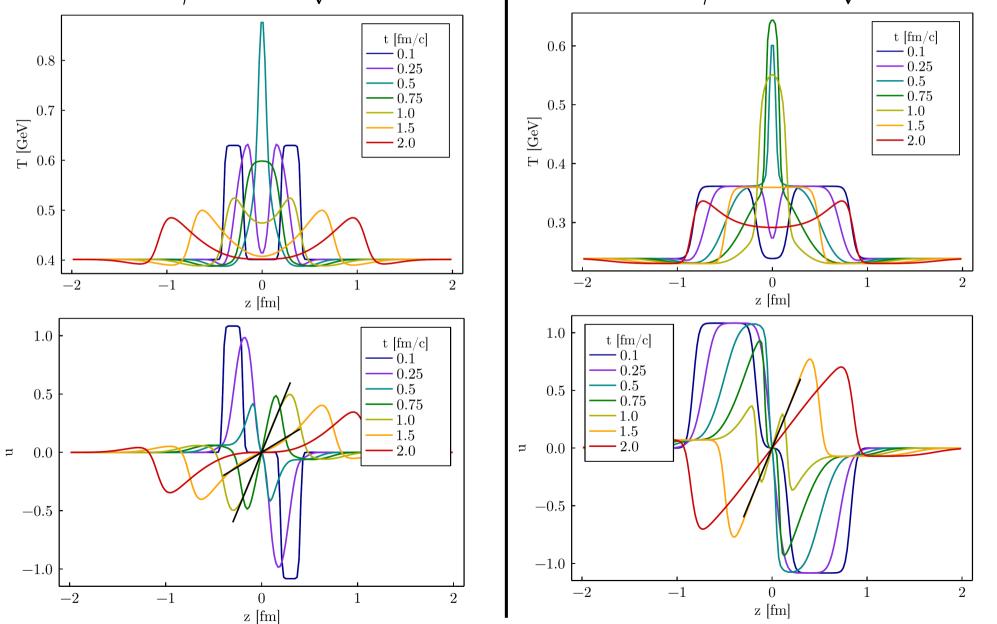
Huge initial pressure gradient  $\rightarrow$  longitudinal expansion

#### First results



#### First results

 $v = 0.9999 \Rightarrow \gamma \approx 71 \Rightarrow \sqrt{s} \approx 133 \text{ GeV}$   $v = 0.999 \Rightarrow \gamma \approx 22 \Rightarrow \sqrt{s} \approx 42 \text{ GeV}$ 

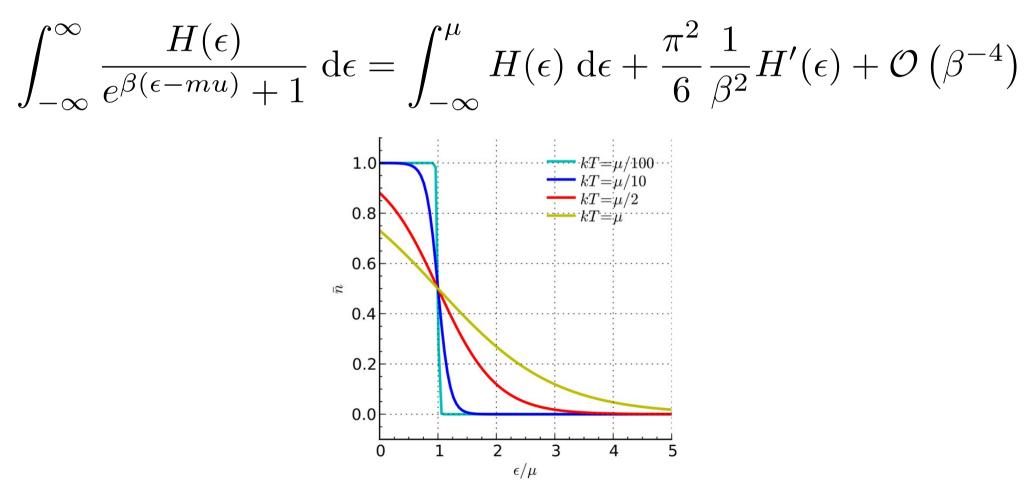


# Summary & Outlook

- Applied hydro to time before collision
  - $\rightarrow$  Heating up
  - $\rightarrow$  Emergent Bjorken flow
- Outlook:
  - Improve initial conditions  $T^{\mu\nu} = T^{\mu\nu}_{in} T^{\mu\nu}_{out}$
  - Initialize viscous fields
    - $\rightarrow$  Examine causality
  - Apply results to traditional hydro

## Sommerfeld Expansion

- At T = 0: Fermions occupy Fermi-sphere
  - $\rightarrow$  Fermi distribution becomes step function
- Including first correction in temperature yields



# Validity of hydrodynamics

- Hydrodynamics assumes local thermal equilibrium
- Deviations from equilibrium

 $\rightarrow$  Dissipative corrections

- Navier-Stokes-type equation  $\pi_{\mathrm{Bulk}} = -\zeta \nabla_{\mu} u^{\mu}$ 
  - $\rightarrow$  Bulk pressure reacts instantly
  - $\rightarrow$  Out-of-equilibrium effects get resolved quickly
  - $\rightarrow$  Causality can be violated
- Add relaxation time  $\tau_{\text{Bulk}} u_{\mu} \partial^{\mu} \pi_{\text{Bulk}} + \pi_{\text{Bulk}} = -\zeta \nabla_{\mu} u^{\mu}$

 $\rightarrow$  System can stay out of equilibrium for longer

# Landau matching

- Decompose  $T^{\mu\nu}$  tensor with respect to vector  $\rightarrow$  Logical choice: fluid velocity  $u^{\mu}$
- Decompose in orthogonal & parallel to  $u^{\mu}$
- Have trace and traceless part

$$T^{\mu\nu} = \epsilon u^{\mu} u^{\nu} + (p + \pi_{\text{Bulk}}) \Delta^{\mu\nu} + \pi^{\mu\nu}$$

## Numeric scheme

- Finite difference method based on Jour. Of Comp. Phys. 230 (2011) 232-244
- Time Derivative given by

$$\phi_{j}^{n+1} = \phi_{j}^{n} - \Delta t \left\{ H \left( \frac{\phi_{j+1}^{n} - \phi_{j-1}^{n}}{2\Delta x} \right) - \frac{\alpha}{2} \left( \frac{\phi_{j+1}^{n} - 2\phi_{j}^{n} + \phi_{j-1}^{n}}{\Delta x} \right) \right\}$$
First order time derivative Numeric viscosity

• Julia implementation with Tsit5() solver